

Planck mass	$\sqrt{\hbar c/G_N}$	$1.22090(9) \times 10^{19} \text{ GeV}/c^2$	[3]
Planck length	$\sqrt{\hbar G_N/c^3}$	$1.61624(12) \times 10^{-35} \text{ m}$	[3]
Hubble length	$c/H_0$	$\sim 1.2 \times 10^{26} \text{ m}$ $= 2.17645(16) \times 10^{-8} \text{ Gyr}$	[43]
parsec (1 AU/1 arc sec)	pc	$3.085\,677\,580\,7(4) \times 10^{16} \text{ m} = 3.262 \dots \text{ly}$	[10]
light year (deprecated unit)	ly	$0.306\,6 \dots \text{pc} = 0.946\,1 \dots \times 10^{16} \text{ m}$	
Schwarzschild radius of the Sun	$2G_N M_\odot/c^2$	$2.953\,250\,08 \text{ km}$	[11]
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solar mass	$M_\odot$	$1.988\,44(30) \times 10^{30} \text{ kg}$	[12]
solar equatorial radius	$R_\odot$	$6.961 \times 10^8 \text{ m}$	[8]
solar luminosity	$L_\odot$	$(3.846 \pm 0.008) \times 10^{26} \text{ W}$	[13]
Schwarzschild radius of the Earth	$2G_N M_\oplus/c^2$	$8.870\,056\,22 \text{ mm}$	[14]
Earth mass	$M_\oplus$	$5.972\,3(9) \times 10^{24} \text{ kg}$	[15]
Earth mean equatorial radius	$R_\oplus$	$6.378\,140 \times 10^6 \text{ m}$	[8]
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luminosity conversion	$L$	$3.02 \times 10^{28} \times 10^{-0.4 M_{\text{bol}}} \text{ W}$ ( $M_{\text{bol}}$ = absolute bolometric magnitude = bolometric magnitude at 10 pc)	[16]
flux conversion	$\mathcal{F}$	$2.52 \times 10^{-8} \times 10^{-0.4 m_{\text{bol}}} \text{ W m}^{-2}$ ( $m_{\text{bol}}$ = apparent bolometric magnitude)	from ab
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$v_\odot$ around center of Galaxy	$\Theta_\odot$	$220(20) \text{ km s}^{-1}$	[17]
solar distance from galactic center	$R_\odot$	$8.0(5) \text{ kpc}$	[18]
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local disk density	$\rho_{\text{disk}}$	$3\text{--}12 \times 10^{-24} \text{ g cm}^{-3} \approx 2\text{--}7 \text{ GeV}/c^2 \text{ cm}^{-3}$	[22]
local halo density	$\rho_{\text{halo}}$	$2\text{--}13 \times 10^{-25} \text{ g cm}^{-3} \approx 0.1\text{--}0.7 \text{ GeV}/c^2 \text{ cm}^{-3}$	[23]
present day Hubble expansion rate	$H_0$	$100 h \text{ km s}^{-1} \text{ Mpc}^{-1}$ $= h \times (9.778\,13 \text{ Gyr})^{-1}$	[19]
present day normalized Hubble expansion rate	$h$	$0.71^{+0.04}_{-0.03}$	[42]
critical density of the universe	$\rho_c = 3H_0^2/8\pi G_N$	$2.775\,366\,27 \times 10^{11} h^2 M_\odot \text{ Mpc}^{-3}$ $= 1.878\,37(28) \times 10^{-29} h^2 \text{ g cm}^{-3}$ $= 1.053\,69(16) \times 10^{-5} h^2 \text{ GeV cm}^{-3}$	derived
pressureless matter density of the universe	$\Omega_m \equiv \rho_m/\rho_c$	$0.135^{+0.008}_{-0.009}/h^2 = 0.27 \pm 0.004$	[42]
baryon density of the universe	$\Omega_b \equiv \rho_b/\rho_c$	$0.0224 \pm 0.0009/h^2 = 0.044 \pm 0.004$	[42]
dark matter density of the universe	$\Omega_{DM} \equiv \Omega_m - \Omega_b$	$0.113^{+0.008}_{-0.009}/h^2 = 0.22 \pm 0.04$	[44]
radiation density of the universe	$\Omega_\gamma = \rho_\gamma - \rho_c$	$(2.471 \pm 0.004) \times 10^{-5}/h^2 = (4.9 \pm 0.5) \times 10^{-5}$	[46]
neutrino density of the universe	$\Omega_\nu$	$< (0.0076/h^2 = 0.015), 95\% \text{ C.L.}$	[42]
dark energy density	$\Omega_\Lambda$	$0.73 \pm 0.04$	[42]
total energy density	$\Omega_{\text{tot}} = \Omega_m + \dots + \Omega_\Lambda$	$1.02 \pm 0.02$	[42]
number density of baryons	$n_b$	$(2.5 \pm 0.1) \times 10^{-7}/\text{cm}^3$	[42]
number density of CMB photons	$n_\gamma$	$410.4 \pm 0.5 \text{ cm}^{-3}$	[47]
baryon-to-photon ratio	$\eta = n_b/n_\gamma$	$(6.1 \pm 0.2) \times 10^{-10}$	derived
scale factor for cosmological constant	$c^2/3H_0^2$	$2.853 \times 10^{51} h^{-2} \text{ m}^2$	
dark energy equation of state	$w$	$< -0.78 \text{ at } 95\% \text{ C.L.}$	[42, 48]
fluctuation amplitude at $8h^{-1} \text{ Mpc}$ scale	$\sigma_8$	$0.84 \pm 0.04$	[42]
scalar spectral index at $k_0 = 0.05 \text{ Mpc}^{-1}$	$n_s$	$0.93 \pm 0.03$	[42]

Quantity	Symbol, equation	Value	Reference, foot
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- Printing Office, Washington, and Her Majesty's Stationary Office, London (2003).
9. JPL Planetary Ephemerides, E. Myles Standish, Jr., private communication (1989).
  10. 1 AU divided by  $\pi/648\,000$ ; quoted error is from the Planetary Ephemerides value of the AU [9].
  11. Product of  $2/c^2$  and the heliocentric gravitational constant  $G_N$ . The given 9-place accuracy seems consistent with uncertainty defining the earth's orbital parameters.
  12. Obtained from the heliocentric gravitational constant [8]  $G_N$  [3]. The error is the 150 ppm standard deviation of  $G_N$  [3].
  13. 1996 mean total solar irradiance (TSI) =  $1367.5 \pm 2.7$  [3]. The solar luminosity is  $4\pi \times (1 \text{ AU})^2$  times this quantity. This value increased by 0.036% between the minima of solar cycles. The TSI in cycle 22. It was modulated with an amplitude of 0.039% during cycle 21 [32].  
Sackmann *et al.* [33] use  $\text{TSI} = 1370 \pm 2 \text{ W m}^{-2}$ , but comment that the solar luminosity ( $L_\odot = 3.853 \times 10^{26} \text{ J s}^{-1}$ ) has an uncertainty of 1.5%. Their value comes from three 1970s papers, and they comment that the error is based on scatter among the reported values, which is substantially in excess of that expected from the individual quoted errors.  
The conclusion of the 1971 review by Thekaekara and Drummond [34] ( $1353 \pm 1\% \text{ W m}^{-2}$ ) is often quoted [35]. The conversion to luminosity is not given in the Thekaekara and Drummond paper, and we cannot exactly reproduce the luminosity given in Ref. 35.  
Finally, a value based on the 1954 spectral curve of Johnson [36] ( $1395 \pm 1\% \text{ W m}^{-2}$ , or  $L_\odot = 3.92 \times 10^{26} \text{ J s}^{-1}$ ) has been used widely, and may be the basis for the higher value of the solar luminosity and the corresponding lower value of the solar absolute bolometric magnitude (4.72) still common in the literature [16].
  14. Product of  $2/c^2$ , the heliocentric gravitational constant  $G_N$  [3], Ref. 7, and the earth/sun mass ratio, also from Ref. 7. The given 9-place accuracy appears to be consistent with uncertainty in  $G_N$  actually defining the earth's orbital parameters.
  15. Obtained from the geocentric gravitational constant [8]  $G_N$  [3]. The error is the 150 ppm standard deviation of  $G_N$  [3].
  16. E.W. Kolb and M.S. Turner, *The Early Universe*, Addison-Wesley (1990).
  17. F.J. Kerr and D. Lynden-Bell, *Mon. Not. R. Astr. Soc.* **22**, 1038 (1985). "On the basis of this review these [ $R_\odot = 8.5 \pm 0.001$  and  $\Theta_\odot = 220 \pm 20 \text{ km s}^{-1}$ ] were adopted by resolution of Commission 33 on 1985 November 21 at Delhi".
  18. M.J. Reid, *Annu. Rev. Astron. Astrophys.* **31**, 345–372 (1993). Note that  $\Theta_\odot$  from the 1985 IAU Commission 33 recommendation is adopted in this review, although the new value for  $R_\odot$  is used.
  19. Conversion using length of tropical year.
  20. M. Fukugita and C.J. Hogan, "Global Cosmological Parameters:  $H_0$ ,  $\Omega_M$ , and  $\Lambda$ ," Sec. 20 of this *Review*.
  21. The final uncertainty arises from dichotomous estimates of the distance to the Large Magellanic Cloud.

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- and  $\Omega_k$ [41]. Thus  $\Omega_{\text{tot}} = 1$  indicates a flat universe.
27. Recent results from both BOOMERANG [37] and MAXIMA-1 [38] indicate  $\Omega_M + \Omega_\Lambda \approx 1$  with  $\approx 10\%$  uncertainties, providing the strongest evidence to date for a flat universe. See discussions elsewhere in this *Review* concerning the remarkable consistency of  $\Omega_M$  and  $\Omega_\Lambda$  measurements by different methods [20,29,39].
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  42. D. N. Spergel *et al.*, *Astrophys. J. Supp.* **148**, 175 (2003), [astro-ph/0302209](#).
  43. Derived from  $H_0$  [42].
  44. Derived from [42].
  45. J. Mather *et al.*, *Astrophys. J.* 512,511(1999). This paper gives  $T_0 = 2.725 \pm 0.002\text{K}$  at 95%CL. We take 0.001 as the one-standard deviation uncertainty.
  46.  $\rho_\gamma = \frac{\pi^2 (k_B T)^4}{15 (hc)^3}$ , using  $T_0$  from Ref. 45.
  47.  $n_\gamma = \frac{2\zeta(3)}{\pi^2} \left(\frac{k_B T}{hc}\right)^3$ , using  $T_0$  from Ref. 45.
  48. Note that one of the priors assumed when deriving this parameter is  $w \geq -1$ .
  49. There are several definitions of  $r$  used in the literature, here  $r$  corresponds to the definition used by Ref. 42.
  50. D. Scott and G.F. Smoot, “Cosmic Microwave Background”, this *Review*.
  51. C.L. Bennett *et al.*, *Astrophys. J. Supp.* **148**, 1 (2003) [astro-ph/0302207](#).

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